

Group Problems #22 - Solutions

Friday, October 21

Problem 1 *Ritz Combination Principle*

Show that the longest wavelength of the Balmer series and the longest two wavelengths of the Lyman series satisfy the *Ritz Combination Principle*.

The wavelengths corresponding to transitions between two energy levels in hydrogen is given by:

$$\lambda = \lambda_{limit} \frac{n^2}{n^2 - n_0^2}, \quad (1)$$

where $n_0 = 1$ for the Lyman series, $n_0 = 2$ for the Balmer series, and $\lambda_{limit} = 91.1$ nm for the Lyman series and $\lambda_{limit} = 364.5$ nm for the Balmer series. Inspection of this equation shows that the longest wavelength corresponds to the transition between $n = n_0$ and $n = n_0 + 1$. Thus, for the Lyman series, the longest two wavelengths are λ_{12} and λ_{13} . For the Balmer series, the longest wavelength is λ_{23} .

The Ritz combination principle states that certain pairs of observed frequencies from the hydrogen spectrum add together to give other observed frequencies. For this particular problem, we then have:

$$\nu_{12} + \nu_{23} = \nu_{13} \implies h(\nu_{12} + \nu_{23}) = h\nu_{13} \quad (2)$$

$$\implies hc \left(\frac{1}{\lambda_{12}} + \frac{1}{\lambda_{23}} \right) = hc \frac{1}{\lambda_{13}} \quad (3)$$

$$\implies \frac{1}{\lambda_{12}} + \frac{1}{\lambda_{23}} = \frac{1}{\lambda_{13}}. \quad (4)$$

Inverting Eqn. 1 above gives:

$$\frac{1}{\lambda} = \frac{1}{\lambda_{limit}} \frac{n^2 - n_0^2}{n^2} = \frac{1}{\lambda_{limit}} \left(1 - \frac{n_0^2}{n^2} \right). \quad (5)$$

To calculate $1/\lambda_{12}$ and $1/\lambda_{13}$, we use the Lyman series numbers and find $1/\lambda_{12} = 3/(4 * 91.1) = 0.00823$ nm⁻¹ and $1/\lambda_{13} = 8/(9 * 91.1) = 0.00976$ nm⁻¹. To compute $1/\lambda_{23}$, we use the Balmer series numbers and find $1/\lambda_{23} = 5/(9 * 364.5) = 0.00152$ nm⁻¹. Computing the left-hand-side of Eqn. 4 gives $1/\lambda_{12} + 1/\lambda_{23} = 0.00823 + 0.00152 = 0.00975$ nm⁻¹, which is equal to $1/\lambda_{13}$ within rounding error.

Problem 2 The Rydberg Constant

When an electron in an atom transitions between orbitals with principle quantum numbers n_1 and n_2 , the emitted (or absorbed) photon has an energy

$$h\nu = |E_{n_1} - E_{n_2}|. \quad (6)$$

- (a) Use Eq. (6) and the expression for the allowed electron energies, E_n , in the Bohr model to show that for *hydrogen* the emitted (or absorbed) photon has a wavelength

$$\lambda = \frac{64\pi^3\epsilon_0^2\hbar^3c}{m_e e^4} \left| \frac{n_1^2 n_2^2}{n_1^2 - n_2^2} \right| \quad (7)$$

$$= \frac{1}{R_\infty} \left| \frac{n_1^2 n_2^2}{n_1^2 - n_2^2} \right|, \quad (8)$$

where R_∞ is called the *Rydberg constant*.

The expression for the allowed energies in hydrogen is given by:

$$E_n = -\frac{1}{n^2} \frac{e^2}{8\pi\epsilon_0 a_0}, \quad \text{with } a_0 = \frac{4\pi\epsilon_0\hbar^2}{me^2} \quad (9)$$

where n is an integer 1 or greater, ϵ_0 is the permittivity of free space, and a_0 is the so-called Bohr radius with m the mass and e the charge of an electron. The energy difference in Eqn. 6 is then given by:

$$h\nu = \frac{hc}{\lambda} = |E_{n_1} - E_{n_2}| = \frac{e^2}{8\pi\epsilon_0 a_0} \left| \frac{1}{n_2^2} - \frac{1}{n_1^2} \right| = \frac{e^2}{8\pi\epsilon_0 a_0} \left| \frac{n_1^2 - n_2^2}{n_1^2 n_2^2} \right|. \quad (10)$$

Inverting this equation and substituting for a_0 gives:

$$\lambda = 8\pi\epsilon_0 a_0 \frac{hc}{e^2} \left| \frac{n_1^2 n_2^2}{n_1^2 - n_2^2} \right| \quad (11)$$

$$= \frac{hc \cdot 8\pi\epsilon_0 \cdot 4\pi\epsilon_0\hbar^2}{e^2 \cdot me^2} \left| \frac{n_1^2 n_2^2}{n_1^2 - n_2^2} \right| \quad (12)$$

$$= \frac{64\pi^3\epsilon_0^2\hbar^3c}{me^4} \left| \frac{n_1^2 n_2^2}{n_1^2 - n_2^2} \right|, \quad (13)$$

where we have used the identity $h = 2\pi\hbar$.

- (b) What is the numerical value of $1/R_\infty$ (don't forget the units!)? You may want to use the *fine structure constant*, α in your calculation:

$$\alpha = \frac{e^2}{4\pi\epsilon_0\hbar c} = \frac{1}{137}. \quad (14)$$

First of all, we see from Eq. 8 above that $1/R_\infty$ must have units of length. Comparing Eq. 8 and Eq. 13 gives:

$$\frac{1}{R_\infty} = \frac{64\pi^3\epsilon_0^2\hbar^3c}{me^4} = \frac{16\pi^2\epsilon_0^2\hbar^2c^2}{e^4} \frac{4\pi\hbar c}{mc^2} \quad (15)$$

$$= \frac{1}{\alpha^2} \frac{2\hbar c}{mc^2} = (137)^2 \left(\frac{2 \cdot 1240 \text{ eV nm}}{511 \times 10^3 \text{ eV}} \right) \quad (16)$$

$$= 91.09 \text{ nm}. \quad (17)$$

- (c) Find the wavelengths of the transitions from $n_1 = 3$ to $n_2 = 2$ and from $n_1 = 4$ to $n_2 = 2$. To which series do these two transitions belong?

Right away we see that the transition $n_1 = 3 \rightarrow n_2 = 2$ belongs to the Balmer series with $n_2 = n_0 = 2$. The transition $n_1 = 4 \rightarrow n_2 = 2$ also belongs to the Balmer series for the same reason. Using Eq. 8 and our value for $1/R_\infty$ gives:

$$\lambda_{23} = 91.09 \left| \frac{9 \cdot 4}{9 - 4} \right| = 656 \text{ nm} \quad (18)$$

$$\lambda_{24} = 91.09 \left| \frac{16 \cdot 4}{16 - 4} \right| = 486 \text{ nm}. \quad (19)$$

As a final note, we compare the phenomenological expression for the transition wavelengths, Eq. 1, and the expression derived from the Bohr model, Eq. 7. When $n_2 = 1$, then Eq. 7 becomes:

$$\lambda_{1,n_1} = 91.09 \left| \frac{n_1^2}{n_1^2 - 1^2} \right|, \quad (20)$$

where n_1 is an integer larger than 1. This expression is precisely that for the transition wavelengths in the Lyman series, where $\lambda_{limit} = 91.09 \text{ nm}$ and $n_0 = 1$. Similarly, when $n_2 = 2$, we have:

$$\lambda_{2,n_1} = 91.09 \left| \frac{n_1^2 \cdot 4}{n_1^2 - 2^2} \right| = 364.4 \left| \frac{n_1^2}{n_1^2 - 2^2} \right|, \quad (21)$$

where n_1 is now an integer larger than 2. This expression is exactly that for the transition wavelengths in the Balmer series, where $\lambda_{limit} = 364.4 \text{ nm}$ and $n_0 = 2$.