

# Modern Physics

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  - Causal structure
  - Lorentz invariance

## (1,3) spacetime

- Primary unit in special relativity  $\Rightarrow$  event
- *Arena*  $\forall$  events in universe  $\Rightarrow$  Minkowski spacetime
- Events take place in a four dimensional structure that contains 3-dimensional Euclidean space + 1 time dimension
- Interval

$$\Delta s^2 = c\Delta t^2 - \Delta x^2 - \Delta y^2 - \Delta z^2 \quad (1)$$

invariant measure for Lorentz-Poincaré transformations



## Lorentz boosts + rotations

- Galilean transformations preserve the usual Pythagorean distance

$$\Delta \vec{r}^2 = \Delta x^2 + \Delta y^2 + \Delta z^2$$

which is invariant under rotations and spatial translations

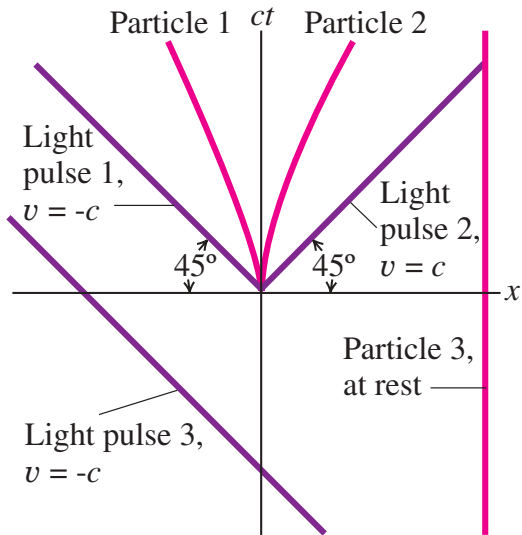
- Lorentz transformations (boosts + rotations) preserve interval

$$\Delta s^2 = c^2 \Delta t^2 - \Delta x^2 - \Delta y^2 - \Delta z^2$$

- Since interval is defined by differences in coordinates  
it is also invariant under translations in spacetime  
(Poincare invariance)
- $(1,3)$  spacetime  $\equiv (1,1)$  spacetime  $\oplus$  rotations

- Trajectory  $\Rightarrow$  connected set of events representing places + times through which particle moves
- Worldlines  $\Rightarrow$  trajectories of massive particles and observers
- At any event on trajectory  $\Rightarrow$  slope is inverse of velocity relative to inertial observer that has  $x = 0$  straight up and with perpendicular set of simultaneity lines of constant  $ct$
- We use  $ct$  rather than  $t$  so that both scales can have same unit
- Particle path forms worldline as particle moves in 1-dimension
- At any point  $\Rightarrow$  slope of worldline is  $d(ct)/dx = (cdt)/(vdt) = c/v$
- Light pulse with  $\pm c$  speed has slope of  $\pm 1$   
giving angles of  $45^\circ$  with the  $\pm x$ -axes
- Since massive particles have speeds less than  $c$   
all worldlines are steeper than those  $45^\circ$  angle
- Nothing known has worldline with slope between  $-1$  and  $1$
- Worldline of particle at rest is vertical and so has infinite slope

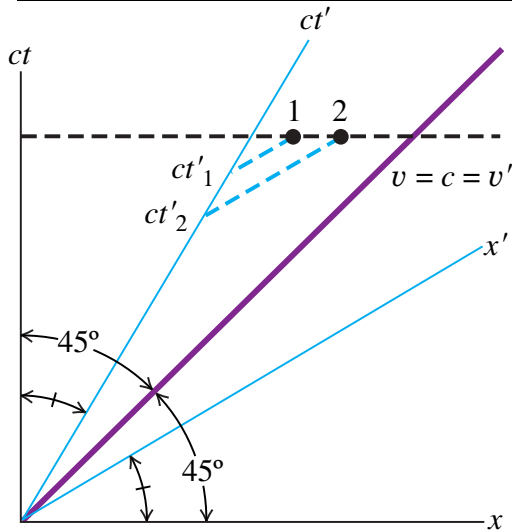
## 3 worldlines and 3 lightlines



## How does the $S'$ reference frame appear on a $ct$ - $x$ spacetime diagram?

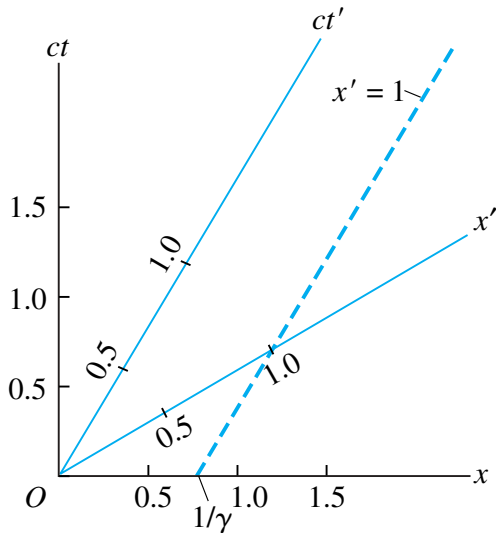
- Recall we always set  $x' = 0$  at  $x = 0$  when  $t' = 0 = t$   
and let  $S'$  move at speed  $v$  in  $+x$ -direction
- $x' = 0$  all along  $ct'$ -axis  $\Rightarrow$   $ct'$ -axis have worldline of slope  $c/v$
- E.g.  $v = 0.600c \Rightarrow ct'$ -axis is at  $\tan^{-1}(1/0.600) = 59.0^\circ$  from  $x$ -axis
- $x'$ -axis is not drawn perpendicular to the  $ct'$ -axis
- Since  $ct' = 0$  all along  $x'$ -axis  $\Rightarrow$  use Lorentz transformation for  $t'$   
 $(t - vx/c^2) = 0$  or  $ct = (v/c)x$  for  $x'$ -axis
- $x'$ -axis is drawn with slope of  $v/c$  on  $ct$ - $x$  spacetime diagram
- E.g.  $v = 0.600c \Rightarrow x'$ -axis is at  $\tan^{-1}(0.600) = 31.0^\circ$  from  $x$ -axis
- $x'$ -axis makes same angle with  $x$ -axis  
as  $ct'$ -axis makes with  $ct$ -axis

## Simultaneity on a spacetime diagram



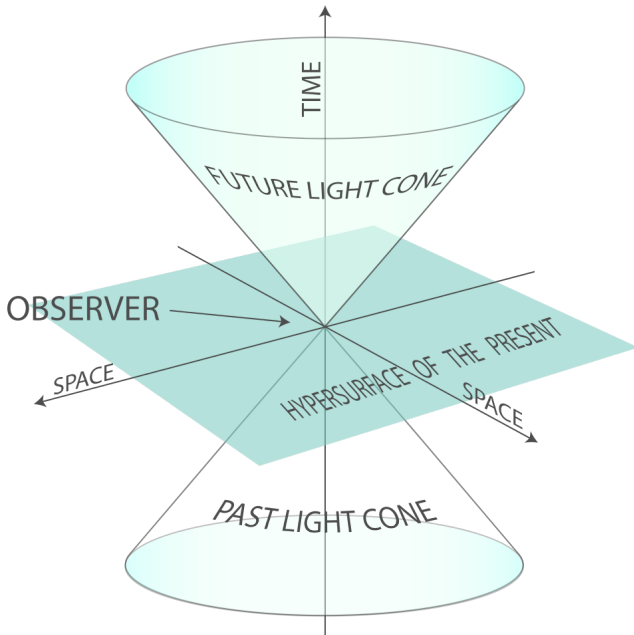


## Length contraction on a spacetime diagram



## Future, past, and elsewhere

- Spacetime around any one event is divided into regions separated by trajectories of light rays emanating from that event
- Separation of events is same for all Lorentz observers since light rays are unchanged by Lorentz transformations
- All events in upper light cone are future of event in question
- From origin event and any event in future  $\Rightarrow$  there exists inertial observer for whom interval between events is a pure time  $(\tau, \vec{0})$  and time of other event is after the now of our original event  $\tau > 0$ .
- Events in backward light cone from original event are in the past
- There exists inertial observer for whom second event is pure time  $(\tau, \vec{0}) \Rightarrow$  but in this case  $\tau < 0$
- All future and past events relative to original event with intervals in any inertial coordinate system have  $\Delta s^2 > 0$ .
- For any elsewhere event  $\Delta s^2 < 0 \Rightarrow$  there exists Lorentz observer for whom events are separated by spatial interval  $(0, \vec{r})$



## 4-vectors

- Event ⇨ a place and a time
- Set of 4 numbers  $(t, \vec{x})$  specifies event in coordinate system
- Designate coordinates with index  $x^\mu$   

$$x^0 = ct, x^1 = x, x^2 = y, x^3 = z$$
- In this notation ⇨ a Lorentz transformations is expressed by

$$x'^\mu = \sum_{\nu=0}^3 \Lambda^\mu{}_\nu x^\nu \quad (2)$$

$$\Lambda^0{}_0 = \gamma \quad \Lambda^i{}_0 = \gamma v^i/c \quad \Lambda^i{}_j = \delta^i_j + (\gamma - 1) \frac{v^i v_j}{v^2} \quad \Lambda^0{}_j = \gamma v_j/c$$

- Einstein summation convention:  
 eliminates summation symbol if same index appears up and down

$$x'^\mu = \Lambda^\mu{}_\nu x^\nu \quad (3)$$

## Spacetime interval

- Given two events  $\Rightarrow$  there is 4 vector

$$\Delta x^\mu = (c(t_2 - t_1), (x_2 - x_1), (y_2 - y_1), (z_2 - z_1)) \quad (4)$$

- Invariant interval is now expressed by

$$\begin{aligned} \Delta s^2 &= g_{\mu\nu} \Delta x^\mu \Delta x^\nu \\ &= c^2 \Delta t^2 - \Delta x^2 - \Delta y^2 - \Delta z^2 \end{aligned} \quad (5)$$

- Metric of Minkowski spacetime:

$$g_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \quad (6)$$

## Metric tensor

- Metric arises directly from physics of spacetime
- It can be used to lower indices  $\Leftrightarrow x_\mu = g_{\mu\nu} x^\nu$
- Inverse metric  $g^{\mu\nu}$  is defined so that  $g^{\mu\nu} g_{\nu\zeta} = \delta^\mu_\zeta$
- Inverse metric raises indices  $\Leftrightarrow$

$$g^{\mu\nu} x_\nu = g^{\mu\nu} (g_{\nu\zeta} x^\zeta) = (g^{\mu\nu} g_{\nu\zeta}) x^\zeta = \delta^\mu_\zeta x^\zeta = x^\mu$$

- Invariance of interval  $\Delta s^2 = \Delta s'^2$  places constraint on form of  $\Lambda^\mu_\nu$

$$g_{\mu\nu} x'^\mu x'^\nu = g_{\mu\nu} \Lambda^\mu_\alpha \Lambda^\nu_\beta x^\alpha x^\beta = g_{\alpha\beta} x^\alpha x^\beta \quad (7)$$

which implies

$$g_{\alpha\beta} = g_{\mu\nu} \Lambda^\mu_\alpha \Lambda^\nu_\beta \quad (8)$$

- (8) can be used to define Lorentz transformation
- Since  $g_{\mu\nu}$  is symmetric there are only ten independent equations
- Only 6 free parameters: 3 to label velocity and 3 to label rotations

## Proper time

- Consider time-like (or light-like) interval  $\Delta s^2 = g_{\mu\nu} \Delta x^\mu \Delta x^\nu$  in any Lorentz frame separating points on particle's trajectory
- Same interval can be expressed in coordinates  $(\Delta t, \Delta x)$  such that at each moment particle is at rest
- Such a frame is called  $\Rightarrow$  instantaneous rest frame
- Since in instantaneous rest frame particle is at rest  $\Delta x = 0$  using interval invariance

$$\Delta s^2 = c^2 \Delta\tau^2 \quad (9)$$

- Because interval is assumed time-like (or light-like) we may take square root of (9) to define proper time interval

$$\Delta\tau = \frac{1}{c} \Delta s \quad (10)$$

## Proper time (cont'd)

- If curved trajectory is time-like  $\Rightarrow$  each segment must be time-like
- Cumulative time is assigned to time-like trajectory  $(t_0, x_0; t_f, x_f)$

$$\begin{aligned} \tau[\text{traj}] = & \sum_{i=0}^{f-1} \left[ (t_{i+1} - t_i)^2 - \frac{1}{c^2} (x_{i+1} - x_i)^2 \right. \\ & \left. - \frac{1}{c^2} (y_{i+1} - y_i)^2 - \frac{1}{c^2} (z_{i+1} - z_i)^2 \right]^{1/2} \end{aligned} \quad (11)$$

- In limit of small segments  $\Rightarrow$  proper time over the trajectory

$$\tau[\text{traj}] = \int_{(t_0, x_0)}^{(t_f, x_f)} \frac{1}{c} ds \quad (12)$$

- Worldline can be specified by  $x^i(t)$  in particular inertial frame
- Alternatively  $\Rightarrow$  worldline specified by  $x^\mu = x^\mu(\tau)$



### 3-velocity vector $\vec{u}$

- Trajectory  $\curvearrowright$  connected set of events  
 coordinatized by some inertial observer as  $(t, \vec{x}(t))$   
 can be parametrized by proper time of trajectory  $\curvearrowright$   $(t(\tau), \vec{x}(\tau))$

$$\begin{aligned}\tau[\text{traj}] &= \int_{(t_0, x_0)}^{(t_f, x_f)} \sqrt{1 - \frac{1}{c^2} \frac{d\vec{x}}{dt} \cdot \frac{d\vec{x}}{dt}} dt \\ &= \int_{(t_0, x_0)}^{(t_f, x_f)} \sqrt{1 - \frac{\vec{u} \cdot \vec{u}}{c^2}} dt\end{aligned}\quad (13)$$

- Elapsed proper time  $\curvearrowright$  functional of the trajectory  
 but function of labels of events at end points of the integral
- Since it is a function of time on trajectory  
 we can derive a differential form of (13)

$$\frac{d\tau}{dt} = \sqrt{1 - \frac{\vec{u} \cdot \vec{u}}{c^2}}\quad (14)$$

## 4-velocity vector $U$

- For a time-like trajectory  $\Rightarrow$  define a four vector velocity

$$U \equiv U^\mu = \frac{dx^\mu}{d\tau} = \frac{dx^\mu/dt}{d\tau/dt} \quad (15)$$

- 4-velocity  $U$  is tangent to world line at each point  
because displacement is given by  $\Delta x^\mu = U^\mu \Delta\tau$

- 4-components of 4-velocity vector  $U$   
can be expressed in terms of 3-velocity  $\vec{u}$

$$\frac{dx^0}{d\tau} = U^0 = c \frac{dt}{d\tau} = \frac{c}{(1 - u^2/c^2)^{1/2}} \quad (16)$$

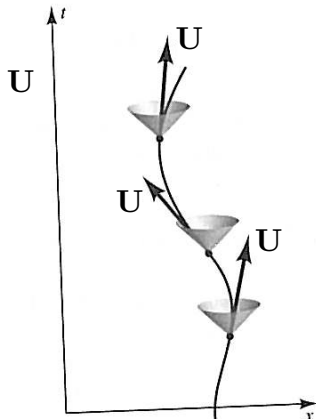
$$\vec{U} = \frac{1}{(1 - u^2/c^2)^{1/2}} \frac{d\vec{x}}{dt} = \frac{1}{(1 - u^2/c^2)^{1/2}} \vec{u} \quad (17)$$

Since  $\Rightarrow$

$$\mathbf{U} \equiv U^\mu = (\gamma c, \gamma \vec{u}) \quad (18)$$

4-velocity is always time-like and future-pointing 4-vector

$$\mathbf{U} \cdot \mathbf{U} = g_{\mu\nu} U^\mu U^\nu = g_{\mu\nu} \frac{(dx^\mu / dt)(dx^\nu / dt)}{(d\tau / dt)^2} = c^2 \quad (19)$$



- @ any point along curve  $x^\mu(\tau)$
- $U^\mu$  time-like tangent 4-vector
- $\mathbf{U}$  lies inside lightcone of point

## Relation between the 4-acceleration vector

$$A = \frac{d\mathbf{U}}{d\tau} \equiv \frac{d^2x^\mu}{d\tau^2} \quad (20)$$

and 3-acceleration vector

$$\vec{a} = \frac{d^2x^i}{dt^2} \quad (21)$$

is more complicated

$$\begin{aligned} A &= \frac{d\mathbf{U}}{d\tau} = \gamma \frac{d\mathbf{U}}{dt} = \gamma \frac{d}{dt}(\gamma c, \gamma \vec{u}) \\ &= \gamma \left( \frac{d\gamma}{dt} c, \frac{d\gamma}{dt} \vec{u} + \gamma \vec{a} \right) \end{aligned} \quad (22)$$

- Note that since

$$\frac{d\gamma}{dt} = \frac{\vec{u} \cdot d\vec{u}/dt}{c^2 (1 - u^2/c^2)^{3/2}} \quad (23)$$

in instantaneous rest frame of particle ( $\vec{u} = 0$ )  $\Rightarrow \mathbf{A} = (0, \vec{a})$

- $A = 0 \Leftrightarrow |\vec{a}| = \alpha = 0$
- For  $A^2$  being an invariant  $\Rightarrow$  evaluate it in rest frame

$$\mathbf{A} \cdot \mathbf{A} = -\alpha^2 \quad (24)$$

- From (24) we see that  $A$  is space-like vector
- By same artifice  $\Rightarrow$  from (18) and (22)

$$\mathbf{U} \cdot \mathbf{A} = 0 \quad (25)$$

4-acceleration vector is always orthogonal to 4-velocity

## Thought-provoking observation

Consider Lorentz transformation  $\Rightarrow$  new frame (prime coordinates) moves with velocity  $v$  along  $z$  axis of original frame

(unprimed coordinates)

Take  $\Rightarrow \cosh(\vartheta) = (1 - v^2/c^2)^{-1/2}$

$$ct' = \cosh(\vartheta) ct - \sinh(\vartheta) z$$

$$z' = -\sinh(\vartheta) ct + \cosh(\vartheta) z$$

$$x' = x$$

$$y' = y$$

(26)

Because  $\cos(i\vartheta) = \cosh(\vartheta)$  and  $\sin(i\vartheta) = \sinh(\vartheta)$

Lorentz boost may be regarded as rotation

through imaginary angle  $i\vartheta$  in  $itc$ - $z$  plane